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Project Report

ETS-19

Magnitudes of Stars on the S-20 System

J. M. Sorvari

14 September 1977

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Lincoln Laboratory

MASSACHUSETTS INSTITUTE OF TECHNOLOGY

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This technical report has been reviewed and is approved for publication.

FOR THE COMMANDER

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MASSACHUSETTS INSTITUTE OF TECHNOLOGY LINCOLN LABORATORY

MAGNITUDES OF STARS ON THE S-20 SYSTEM

J. M. SORVARI Group 94

PROJECT REPORT ETS-19

14 SEPTEMBER 1977

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ABSTRACT

An S-20 magnitude system is defined, and a relation between it and B,V-photometry is established. A catalog of $\rm m_{_{\rm V}},$ (B-V), and $\rm m_{20}$ for 323 stars is given.

I. INTRODUCTION

Several institutions, including Lincoln Laboratory, are currently involved in observational projects which utilize image intensifier tubes for the detection and measurement of point sources. A variety of instruments, such as vidicons and image disectors, follows the intensifier tubes, but all share the spectral sensitivity of the first photo-sensitive surface of the intensifier, usually the multialkali type designated S-20.

Supplementary to the imaging system, it is desirable to have a classical astronomical photometric system with a set of known stars which serve as the system's intensity standards. Such a system can aid in the calibration of the imaging system measurements, correction for the effects of the earth's atmosphere and evaluation of the effects of the sky background. The simplest system for accomplishing this is one having the same spectral sensitivity as the imaging system, in this case the unfiltered S-20 response. This Report presents observations made on such a system -- the 6"-photometer at the GEODSS ETS.

The purpose of this catalog is use by the ETS automatic extinction package.* Additional uses include use as a reference system for variable star observers or for comet watchers. The automatic extinction package provides real-time measurement of

^{*}J. M. Sorvari and C. E. Beane, "Automatic Real-Time Extinction Measurement," Project Report ETS-17, Lincoln Laboratory, M.I.T. (12 September 1977).

atmospheric extinction and night sky brightness; it requires a sequence of reference stars distributed over the whole sky.

Since no advantage accrues from having regions of star density much greater than the mean, and since such concentrations do take additional computer resources, a significant fraction of the candidate stars was rejected in order to smooth the distribution. The somewhat higher density of stars near the equator reflects the greater importance of extinction measurements in this "synchronous band" of the sky. Limits on the magnitudes included were chosen to provide sufficient accuracy, avoid the need for changing the measuring range of the electronics and to allow both star and sky measurements to be made with the same field stop (~2').

II. THE m20 SYSTEM

Observations at the ETS were made with an uncooled EMI 9785B photomultiplier tube, having an S-20 spectral response. The response to starlight was measured as anode current using a Pacific Photometric Industries Model 124 photometer. Magnitudes calculated from these measurements are designated by the symbol m_{20} .

The standard stars for the m₂₀ system are listed in Table I. Observations of standard stars were made on nights of "photometric quality" over a wide range of zenith distance. This allowed mean corrections for atmospheric extinction to be applied to both standards and unknowns. Although nights of genuinely high photometric quality are very rare at the ETS, it was found possible to satisfactorily reduce most night's data by making two to three times the usual number of standard observations. Each night's data was then brought to the standard system through comparison of the standards. In each case a simple additive constant (i.e., a zero-point drift) was found to be sufficient. The mean error of a single observation was <0.000.

A total of 21 stars, in addition to the standards, with known Johnson B and V magnitudes was observed. These stars had (B-V) colors ranging from 0.40 to 0.91. A linear least squares fit between the systems yields the relation

$$m_{20} = m_{V} - .38 + .61(B-V)$$

The value .38 is needed to fit the normalization at (B-V) = .62.

This was used instead of the usual astronomical normalization at (B-V) = .00 for two reasons. First, no stars as blue as (B-V) = .00 were observed, and second, because of the anticipated use of the $m_{2\,0}$ system at ETS, it is convenient to have m_V and $m_{2\,0}$ approximately equal for nearly solar colored objects. The mean deviation from the linear fit is 0.03 and is apparently due entirely to observational uncertainty in $m_{2\,0}$, m_V and (B-V).

TABLE I STANDARD STARS FOR THE ETS M_{20} SYSTEM

	STAR		m ₂₀
]	HD 10307	BS 483	4 ^m 94
	20630	κ Cet	4.87
	34411	λ Aur	4.71
	65228	11 Pup	4.25
	86728	20 LMi	5.41
	111812	31 Com	4.94
	142373	χ Her	4.59
	157214	72 Her	5.38
	182835	v Aql	4.67
	193370	35 Cyg	5.21
	220657	U Peg	4.40

III. THE ETS m20 CATALOG

Ordinarily a catalog of magnitudes would be built up through observation of a large number of stars on several photometric nights each. However, a great saving in observing time may be realized at only a moderate cost in precision by applying the results of Section II to stars with B,V-photometry. This has been done to give the catalog of 323 stars listed in Table II.

In order that this list be most useful to the work at the ETS, the stars have been limited primarily to the following ranges: $\delta > -25^{\circ}$, .4 <(B-V)<.9, $3.5 < m_{20} < 6.0$. An attempt has been made to avoid star density varying strongly with right ascension. Of the 323 stars, 121 have $\delta > +10^{\circ}$, thus the mean separation of these stars is about 12° . For $-20^{\circ} < \delta < +10^{\circ}$, there are 175 stars for a mean separation of 8° .

The first two columns of Table II give the stars numbers from the General Catalog (GC) and Harvard Annals (HD). These are followed by right ascension (RA), annual variation in RA, declination (DEC), and annual variation in DEC, for 1978.0. Next come the Johnson V-magnitude (m_V), the color (B-V) and the S-20 magnitude (m_{20}). The last column contains a code with the following meanings:

- 0 Star not observed at ETS; $m_{2\,0}$ is constructed from $m_{_{\rm V}}$ and (B-V) as in Sec. II
- 1 Star observed 1-3 times; m20 is mean of

observations and constructed magnitude

- 2 Star observed ≥4 times; m₂₀ is mean of observations
- 3 Standard star; $m_{2\,0}$ is standard value The estimated precision in $m_{2\,0}$ ranges from better than 0.01 for the standard stars to 0.03 for the constructed magnitudes.

Many of the stars listed are components of binary systems. In these cases the entire system is the intended object, and the numbers listed are for the combined light.

TABLE II

gC	HD		RA	K	ANV		DEC	D	ANV	m N	B-V	M20	O
3334	224930	00 h	01m	8.00	3.1	9	1	. 28		m7	9	m	C
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2	67		6	7		37	20	48	20	5.15	. 43	5.03	00
486	11	, 4	4	9		01				1	00	6	C
9	19		4	9		-03				7	1		0 0
0	22		4	4.		00-				6	4		0 0
759	3546	7	37 2	23.2	3.2	29	11	32		~	∞	5	0
1	9	* 17	00	2.		21		52	19	5.87	. 85	6.01	0
959	9	7	1	3		0			19		∞	6	0
	4676	7	47 4	10.6			49	19		0.	5	0	0
0	81	7	6	-		-				-	5	-	C
0	11		6	5		9				3	4	~	0
	92	010	6	3	3.4	41	57	53	19	5.65	09.	5.64	0
0	47	1	3	7	0	0				9.	4	5	0
0	03		∞	0	0	0				00	9	00	0
0	92	7	00	0		5				6	9	0	0
817	9021	(4	29 3	33.6	4.5	70	60	08	19	00	4	7	0
∞	2	(*)	2	9	0				18	5.74	. 65	5.76	0
5	030	4	0	9			30			6.	9	0	~
∞	047	4	-	7		2	-			2.	00	3	0
N	070	4	3	2						5	-	5	0
191		4	46 0	6.20	4.3	63	44	42		9.	∞		0
-	223	r)	6	0	.0		-		17	5.87	.62	5.87	0

	Ø	C	0 0	0 0	0 0	0			0 0	0	0	0	0	0	0	0	C)	0	0	0	0	C	0) M	0
	M2 0	m	2	. 10	000	5.47	5	9	000	0	5.18	9	0	· ·		5.71	9	000	0	~	5.87	9	5	0	000	6.04
	B-V	_00	9	5	4	. 58	∞	4	9	9	09.	4	∞	00	4	. 52	5		4	4	. 48	4	4	5	9	. 86
	m A	E O	9	9	0	5.50	0	-	00	5	-	1	00	-	∞	5.77		0	-	4	5.96	00	1	0	00	5.90
continued	ANV	1	-	-		17				17						15	15				15					13
TABLE II (co	DEC	00026'4	58 19 1	08 28 0	-10 09 14	02 29 4	8 44 4	5 40 5	4 07 3	01 39 17	3 54 5	5 20 2	05 29 5	2 43 2	11 57 5	09 3	00 47		49 08 1	8 39 5	08 17 4	20 34 5	5 21 4	01 16 3	03 17 27	6 59 3
	ANV	3.3								3.1	0		3.2					3.8		0			0		3.2	
	RA	2m40.5	7 07.	0 11.	17	1 39.	6	4 30.	5 42.	6 52	1 32.	2	4 43.	5 53.	8 29.	0 6	90 0	40 51.1	2 41.	4 04.	5 03.	6 49.	8 37.	1 38.	18 12.3	8 36.
	HD	12641	295	342	345	361	13611	387	397	421	480	15798	919	592	662	199	67	73	689	720	826	840	869	666	20630	190
	9	2488	54	19	62	65	2656	10	73	1	9	3045		11	13	21	3235	24	27	31	5	56	09	83		97

TABLE II (continued)

O	0	0		0	-	0	0	0	0	0	0	0		0	0	0	0	0	0	0	0	0	0	0	0
m20		5	-	00		-	4	~	2	5.95	0.	4.	6	5	5.40	6.	0.	0.	6.	5.49	7.	5	2	4.92	6.
B-V		5	4	∞	.57	4	.68	5	5	9	5	5	∞	∞	. 82	9	9	2	9	.63	9	4	9	.65	9
m V	m.5	.5	~	1	4.28	2	4.46	3	3	0	-	.5	4.83	4.	. 2	1.	∞	2.	6			9.	0.	4.90	6.
ANV		3			11		11				10		6	9	6	6	00	7	9	7	9	9	2	2	2
DEC	805711	8 48 1	5 59 0	9 31 5	7	23 18 4	-03 01 07	00 19 3	2 46 0	1 57 0	0 38 3	7 59 0	09 12 34	7 41 0	65 05 2	6 04 4	5 54 5	02 30 5	-08 50 07	16 58	05 42 3	6 24 3	60 24 4	18 37 01	2 31 0
ANV	3.52				9		3.0				0	4.			5.7					2.7				3.6	
RA	03h23m37.55	8 08	0 54.	53.	5 44.	5 53.	53 11.0	1 29.	3 00.	4 01.	6 13.7	7 08.2	12 44.3	4 15.	8 35	9 30.	7 18.	6 29.	1	6 3	7 31.	8 02.	1 27.	6.80 90	6 23.
HD	112	138	177	204	4	37	24554	545	562	298	500	599	-	969	702	749	830	939	000		05	192	191	32923	309
29	07	14	21	4244	31	54	4706	85	89	91	8	97		13	19	2	43	63	5759	84	98	08	13	6255	26

TABLE II (continued)

O	0	0	~	0	0	0	0	0	0	0	0	0	0	0	Н	0	0	0	0	0	0	0	0	0	0
m ₂₀	5.m01	6.	1.	6.	2.	6.	0.	5.53	2.	00	0.	6.	0.	5.64	. 2	r.	.5	0.	5.69	. 2	0	9.	~	5.63	
B-V	m44	9	9	5	4	4	9	.45	4	4		∞	9	88.	∞	4	5	4	9	. 43	5	1	4	.40	
m ^	5m12	00	1.	6.	3	0.	0.	5.64	4	6.	6.		6	5.48	7.		9.	0.	5.70	3		5	4.	1	1.
ANV	5.	4	3	4	4	4	4	3	2	2	1	٦	-	~	0	0	0	0	-01	0	0	0	0		-05
DEC	04028	01 00 40	40 04 55	0 60	02 34 2	79 12 3	24 47 3	-19 42 49	30 28 5	06 48	1 09 4	09 30 5	4 05 5	14	3 15 5	22 46 0	0 30 0	2 16 4	05 06 31	3 27 5	00 55 5	8 38 2	79 35 3	0 30	25 24 1
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RA	7m385	10 37.	7 35.	7 51.	02	8 53.	0 51.	25 02.6	7 13.	1 49.	5 26.	5 39.	7 29.	9	2 47.	8 52.	4 27.	5 12.	909	0	4 09.	5 25.	2 31.	9 42.	53 57.5
НД	325	364	441	472	34658	356	516	573	726	38089	00	852	885	907	41116	244	331	338	358	43905	506	713	658	993	50692
29	29	36	49	51	6059	45	59	70	00	7151	22	22	28	31	2929	84	00	03	05	8151	29	19	71	95	9064

TABLE II (continued)

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m ₂₀	m	. 17	000		5.15	7	- 00	4	00	5.02	1		2	0	5.48	4	. 7	0	0	5.02	9	1) \	. 10	5.61
B-V	00	15	10	0	.46	5	5	5	9	44	A	5	1	-	. 93	A	00	5	7	.46	5	A	4	" 15	. 61
m A	E S	1	0	0	5.25	7	00	4	6	5.13	6	-	7	2	5.29	5	3	0	0	5.13	5	9	0	2	5.62
ANV	0	C	0	0	-07	0	80-	0	0		0		0		-10		-	-		-12		-	-		-12
DEC	802711	7 16 5	7 11 0	5 43 4	21 29 30	11 30 3	-08 49 54	22 14 5	05 54 4	04 03 3	4 38 1	13 50 2	5 22 1	22 49 1	02 17	13 44 0	7 42 2	17 42 5	33 2	27	20 00 3	03 40 4	9 40 2	06 4	5 05 5
ANV	(V) .				3.6		2.9								3.1	9			2.9					3.2	
RA	5m22.	7 14 12.	18 36	4 46.	6 26.	6 49.	31 02.5	3 06.	5 24.	6 10.	7 44.	0 45.	42	5 54.	7 12.	9 38.	0 28.	0 57.	17 21.3	8 43.	0 22.	3 29.	7 47.	34 40.9	7 15.
HD	052	557	57006	852	872	906	59984	053	080	106	111	409	64235	522	534	814	831	825	69830	989	044	095	176	72945	290
CC	9082	09	7	92	95	97	10090	013	019	021		062	4	075	077	11118	113	114	132	134	139	147	09	11781	8

	Q		0	0	0	0	0	0	0	0	0		0	0	0	0	3	0	0	0	0	0	2	0	0	0
	m20	E	1	4	0	5.77	00	0	0	· ~	5.39		1.	00	0.	4.29	4	9.	7	1	5.26	9	00	0.	-	5.78
	B-V			6	9	. 53	4	9	9	.64	-	1	9	5	9	.93	9	4	3	5	. 94	10	∞	5	4	. 83
	m A	E C	9	3	6	5.83	6.	6	7.		5.40	5	.5	6	0	4.10	~	00	.2	00	5.07		1.	6.	2.	
(continued)	ANV	\leftarrow	H	H	-14	-	\vdash	H	\vdash	-16		-		-16				\vdash	H	-19	-		-	\vdash	-19	
	(E)	0		0	1 00	5	\vdash	-	5		9 12	4	4		7	1 33	5	5	7	~		4	2	4	12	7
E II	DE	203	7 0	3 2		0 9	1 5	6 4	1 5	05 58	0 6	9 5	6 2	3 4	0 9	4 4	2 0	9 3	2 4	0 9	1	2 0	2 0	6 4	0 01	3
TABLE					1		7		Ī		0	0	(*)	-2		-	(1)	7	80	5	0				-2	
	ANV	2.5					0		0	3.0				2.7	0						3.1				2.9	
	RA	10.5	5	0	13.0	2	ď	6	2	41.4	9	4	2	14.8	0	5	4.	5	4.		39.5	9	6	0	24.9	2
					53		59			26		7	7	41	1	0	6	∞	∞	59	\sim	2	_	1	52	2
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	HD	375	439	491	76151	693	4	841	049	18	185	221	263		473	544	672	944	008	90839	161	188	212	221	94388	448
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TABLE II (continued)

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m20	E.	5	2	6	5.19	00	9.	0.	00	5.92	9	.2	1.	5.81	•		2.	6.	∞	3.73	0.	∞	0.	4.99	6.
B-V		5	7	4		-	4	1	3	.95	∞	3	5	. 48	5	5	3	4	. 70		4	4	5	.53	5
m	m.	5	3	0	5.22		1.	0.	6	5.72	5	4.	00	5.90	0.	00	4.	0.	4.80	00		6.	Γ.	5.05	6
ANV	2	-	-19	-		2		-	H	-18			-	-17			-	\vdash	-16	-	-	-	H	-14	
DEC	703913	0 27 55	17 38 4	6 04 4	09 3	18 11 2	5 02 5	13 53 4	37 17 4	-05 17 04	16 04 0	09 12 3	25 11 4	-00 44 34	05 53 4	25 43 0	13 06 1	51 57 0	-02 07 49	05 33 5	5 54 2	4 57 1	7 25 5	01 50 51	0 21 5
ANV	() ·				3.0					3.1		0		3.1					3.1					3.1	
RA	2h50m375	2 35.	08 55.	0 52.	40	7 15.	3 24.	7 21.	3 48.	34 22.4	3 18.	5 32.	9 23.	12 32.4	4 51.	7 45.	8 12.	4 26.	7 03		9 28.	6 20.	4 22.	18 11.1	2 17.
НД	1181	1341	1437	1464	3	1561	1656	1717	1821	118219	1960	2325	2399	124425	2485	2527	2545	2666	126868	2950	3081	3408	3606	136202	3710
00	745	771	783	787	17975	800	813	821	835	18366	856	904	912	19188	924	930	931	946		2	997	034	053	20591	69

5.89 5.54 3.71 4.56 4.59 3.76 5.40 5.40 4.83 4.46 5.38 6.02 4 m 82 5.71 5.52 5.91 4.42 4.05 4.45 5.52 4.67 0 #45 .50 .86 .69 . 922 B-V 58 52 51 51 4.43 5.78 5.38 5.87 4.43 4.63 3.85 5.46 5.41 .53 .92 97048 m V 7.50.01 44044 5 5 m 4 4 44450 -05 -05 -05 -03 -08 -07 -07 -06 -10 -10 -10 -09 112 -10 -11 -11 -10 TABLE II (continued) -10°14'39" -08 43 18 -19 13 53 02 35 05 07 25 14 55 23 42 46 46 17 48 46 05 43 13 07 24 50 39 46 06 57 DEC 18 48 18 07 33 29 34 05 29 36 30 28 37 22 37 16 13 57 44 10 115 115 133 58 -02 01 38 -10 65 54 -21 32 07 -111 -20 -08 -24 -16 m 5 m 9 4 33102 22010 33.33.33.39.59.09.59 ANV H 88 H R H HERRY mm nmo 12322 mmmmm 15^h23^m00.50 37 28.6 37 38.1 42 55.3 45 22.3 54.7 26.1 05.4 11.9 28.3 W L W L 4 90000 40H0W 09. 06. 25. 18. 111. 335. 36. 54. 552. 41. 15. 04 13 19 19 03 06 14 19 29 35 44 2 5 5 5 5 5 5 5 5 17 91 54905 55885 56897 57214 57978 137052 139460 139446 140538 142373 142860 143333 143761 144284 144069 144608 14623 147084 148786 149661 150557 150997 151769 HD 22321 22460 22502 22643 22871 23092 23274 23423 23446 23614 21340 21408 21495 21527 21572 21593 21659 21864 21969 22200 21029 21031 21155 21201 20699 25

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m 2 0	m.	~	1.		9	4	-	2.	9.	4.19	6	-	4.87	0.		-	1.	· .	. 2			1.	5.21	.2	
B-V	m.39	1	∞	2	4	3	3	5	3	. 84	5	0	.38	0		4	9	5	.38			1	.60	5	-
m A	H.	3	5	0.		9.	2	3	1.	4.05	0.	6.	5.02	00	9.	2	1.	7.	5.44	0	6.	.5	5.22	2.	
ANV	0	0	0	-01	0		0	0	00	0			01							04	0 5		0.5		
				01					55				22						60	04			19		
DEC	0			28					10				23						39				52		
	0	0	0	16	0	0	2	2	-08	0			64		0 8	0	5	\vdash	90		50		32		
ANV	(J) •									3.0			0.4						2.9				2.3		
	2.	5	5		6	6	3	0	53.1	0	-	6	46.3	5	5	4.	1	00	34.9	5	0	6	12.0	4.	
RA	25m	6	0	0	52	6	0	\vdash	01	4			13			9	7	~		6			99		
	17h						18																		
HD	5795	6861	5883	163989	291	6425	6466	6458	6476	53	6590	6620	168151	6865	968	8669	7163	7139	171834	446	755	7549	176051	7630)
25	361	370	373	24343	32	450	453	455	456	9	470	472	24916	503	51	517	536	537	25427	84	593	595	26030	607)

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O	00 00	00000	00000	00000	00000
m20	4.79 5.18 4.67 4.79	5.33 4.93 5.05 89	5.09 5.60 3.85 5.85	5.77 5.66 5.73 3.75 5.21	4.63 5.69 4.51 5.11
B-V					.37 .69 .95 .71
m ^	4,85 5.00 4.67 4.69	5.45 5.06 4.86	5.13 5.55 3.71 4.70 5.88	5.78 5.75 5.90 3.56 5.16	4.78 5.65 4.31 5.06
ANV	007	80000	00 00 00 00 00	111111111111111111111111111111111111111	12 12 13 13
DEC	-25 ⁰ 17'47" -05 27 31 00 17 37 69 37 25 50 10 08	-04 41 54 45 28 17 -16 10 39 -19 48 55 33 40 32	10 21 33 24 56 06 06 21 02 -26 21 36 -10 01 02	17 00 32 62 00 41 45 30 39 -12 36 51 34 54 47	-17 53 17 -09 55 47 -01 10 59 10 00 28 -25 21 06
ANV	3.2 3.2 3.1 -0.1	3.2	3.7	2.7 1.0 1.9 3.3	3.5 2.9 6.9
RA	19 ^h 14 ^m 11.5 19 22.3 25 23.6 32 24.2 35 51.1	36 37.1 40 09.4 41 15.6 45 04.8	49 58.3 51 05.9 54 14.0 54 29.5 58 35.5	20 03 06.6 11 13.3 15 17.6 16 50.1 17 48.2	27 36.4 31 11.4 37 12.1 38 03.5 44 47.7
HD	179950 181391 182835 185144	185124 186155 186005 186648	187691 187982 188512 188376 189340	190406 192455 192985 192947 193370	194943 195564 196574 196755
GC	26516 26669 26838 27050	27143 27249 27255 27349 27369	27480 27516 27587 27583 27689	27835 28071 28174 28200 28242	28481 28563 28725 28756 28929

TABLE II (continued)

O.	10000	00000	00000	00010	00100
m20	4.19 4.03 5.88 4.70	5.12 5.56 4.70 4.42 3.58	3.85 6.04 4.45 5.28	4.19 5.20 4.88 4.44	4.92 4.92 4.20 5.90
B-V			00000		
m D	4.72 4.72 4.72	5.22 5.53 4.51 3.73	3.91 5.86 4.28 5.48	4.02 4.72 4.51 4.28	5.07 4.27 5.74 3.65
ANV	13 14 14 14	14 15 15	15 15 15 16	16 16 16 17	17 17 18 18 18
DEC	57°30'00" 16 02 39 -05 42 35 27 00 50 -09 04 02	04 12 28 -05 54 40 -11 27 42 09 55 01 37 57 03	05 09 21 -04 39 16 -16 55 47 -12 58 26 -21 54 14	45 29 37 -14 08 47 -18 58 04 28 38 34 61 01 07	-13 39 24 73 04 29 33 04 12 -05 29 51 -00 08 00
ANV	1.55 2.8 3.2 3.2	22.22	33.3.0	2.3	3.3
RA	20 ^h 44 ^m 48 ^s 4 45 37.3 50 15.7 51 11.1 51 28.2	57 58.5 21 02 54.8 08 23.9 13 24.5 13 54.8	14 43.4 19 54.9 21 01.4 22 59.4 27 28.3	33 09.1 40 21.1 41 25.9 43 09.2 44 48.7	52 05.9 58 56.7 22 09 00.3 15 57.5 27 41.7
HD	198084 197963 198571 198809 198743	199766 200496 201381 202275 202444	202447 203222 203387 203705 204381	205435 206301 206453 206826 207260	207958 209369 210459 211434 213051
CC	28956 28965 29078 29112 29109	29276 29417 29571 29697 29723	29735 29877 29903 29957 30059	30207 30354 30382 30437	30631 30800 31016 31163 31398

TABLE II (continued)

O	0	0	Н	0	0	0	0	0	0	0	0	0	0	3	0	0	0	0	0	0	0	0	0
M _{2 0}			0	7.	0.	00	.5	5	5.57	1.	00	4.	4.	4.40	1.	9.	5.55	0.	6.	9.	4	5.05	00
B-V	m39	4	5	1	4	∞	9	1	.91	9	4	∞	4	.61	0	3	. 44	5	∞	0	9	. 93	4
m A	m.	5.20	۲.	9.	۲.	7.	4.	4.	5.39	9.	0	0.	.5	4.44	.5	00	5.66	-	00	4.	4.	4.86	0.
ANV		18							20					20			20					20	20
DEC	401910	-20 49 19	12 03 3	1 34 4	9 43 0	5 06 1	0 39 0	5 16 0	02 00 28	2 34 3	43 25 3	13 34 4	5 14 4	23 16 57	2 38 2	1 22 0	01 58 48	5 30 2	7 56 1	02 53 0	2 05 3	-03 40 40	6 44 2
ANV	(C)	3,3							3.1					3.0			3.1						3.1
RA	8m51.5	33 29	5 35.	0 30.	1 17.	4 02.	6 22.	7 11.	07 33.2	8 44.	9 25.	7 57.	8 15.	24 16.5	8 02.	3 00.	35 15.8	8 49.	0 37.	6 48.	7 45.	7 32.	58 10.7
HD	1323	213845	1564	1638	1638	1671	1701	1865	218527	1864	1880	1983	1987	220657	2111	2167	221950	2236	2257	325	2338	245	H
CC	142	31516	177	188	189	196	200	223	32233	226	228	246	246	32585	266	277	32818	287	291	302	305	324	33262

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An S-20 magnitude system is defined, and a relation between								
it and B, V-photometry is established. A catalog of m,, (B-V), and								
m ₂₀ for 323 stars is given.								